

North American Nanohertz Observatory for GWs



http://nanograv.org/

Our mission is to detect nHz frequency gravitational waves from super-massive black hole binaries.

We are over 100 students and scientists and welcome new collaborators and members.



Astro2020 Project White Paper

The NANOGrav Program for Gravitational Waves and Fundamental Physics



The North American Nanohertz Observatory for Gravitational Waves

July 10, 2019

Thematic areas: Multi-messenger astronomy and astrophysics; Cosmology and fundamental physics; Formation and evolution of compact objects.

Contact author: Scott Ransom (NANOGrav Chair), NRAO, scott.ransom@nanograv.org

Authors: A. Brazier (Cornell), S. Chatterjee (Cornell), T. Cohen (NMT), J. M. Cordes (Cornell), M. E. DeCesar (Lafayette), P. B. Demorest (NRAO), J. S. Hazboun (UW Bothell), M. T. Lam (WVU, RIT), R. S. Lynch (GBO), M. A. McLaughlin (WVU), S. M. Ransom (NRAO), X. Siemens (OSU, UWM), S. R. Taylor (Caltech/JPL, Vanderbilt), and S. J. Vigeland (UWM) for the NANOGrav Collaboration (~ 50 institutions, 100+ individuals)

Ransom et al. 2019, Astro2020 Whitepaper

But not just NANOGrav...



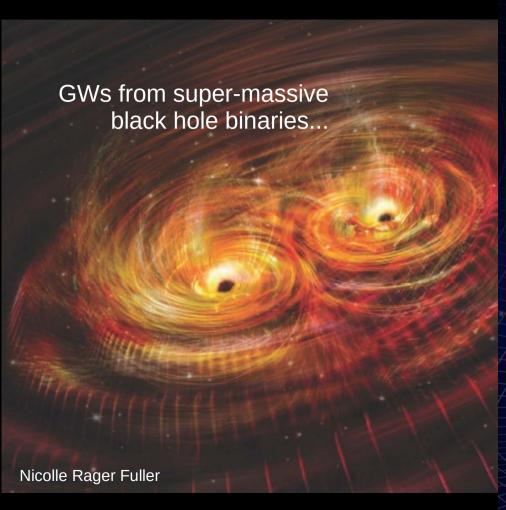


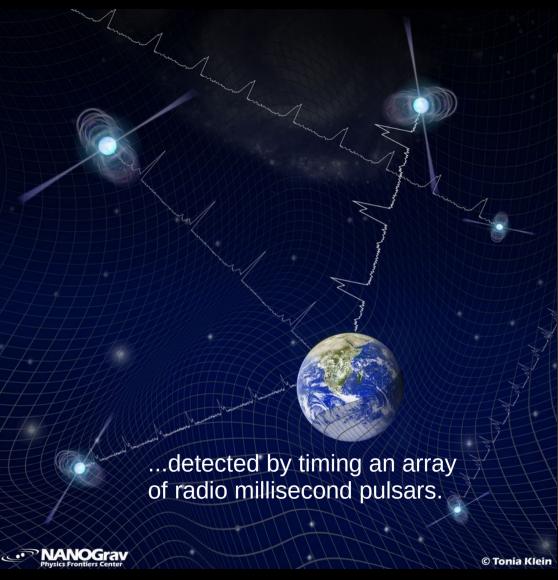
Parkes Pulsar Timing Array





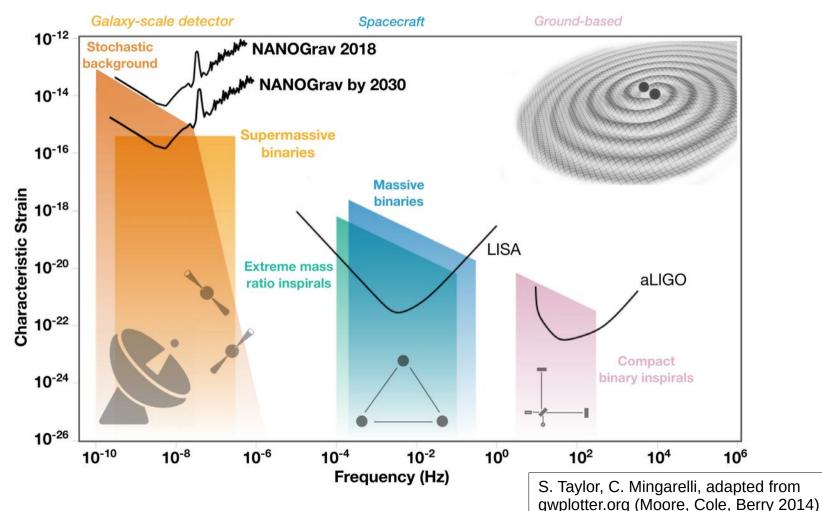
European Pulsar Timing Array







The Gravitational Wave Spectrum













Both Arecibo and the GBT are Crucial



Each telescope provides ~50% of our GW sensitivity

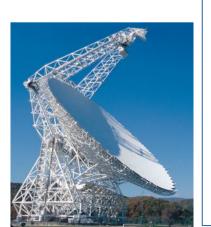
- Arecibo has 4-5x sensitivity
- **GBT** has 3x sky coverage

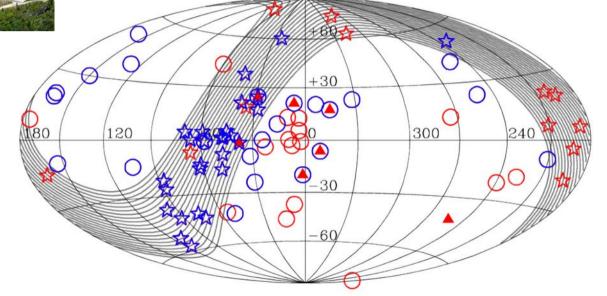












12.5-Year Data Set More Recent Additions



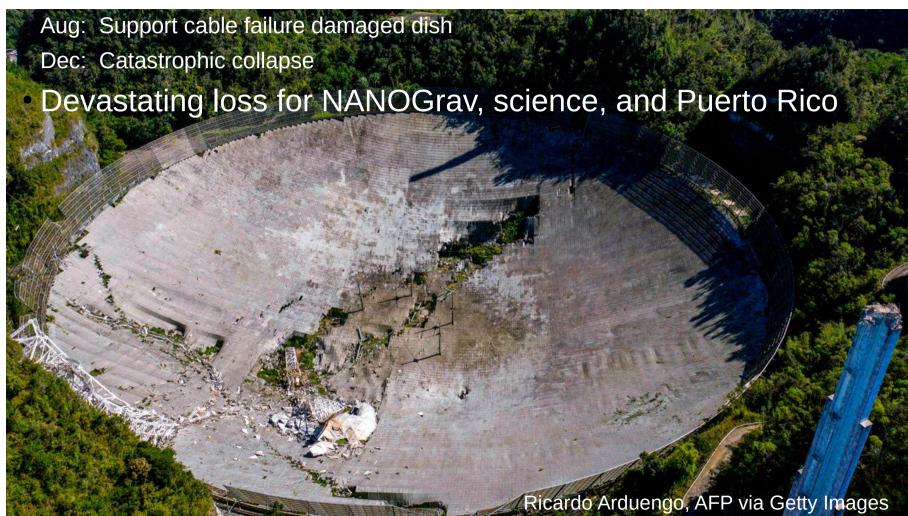








Loss of Arecibo













Arecibo's Pulsar Legacy

1968 Crab spin period (Lovelace et al)

1970 pulsar nulling (Backer)

1974 1st binary pulsar (Hulse & Taylor)

1982 1st millisecond pulsar (Backer et al)

1989 indirect GWs from B1913+16 (Taylor & Weisberg)

1990 1st eclipsing MSP (aka black-widow; Fruchter et al)

1992 pulsar planets (1st exoplanets!; Wolszczan & Frail)

1994 microsec long-term timing (Kaspi, Ryba, & Taylor)

2008 an MSP in an eccentric orbit (Champion et al)

2014 1st non-Parkes FRB (Spitler et al)

2016 1st FRB repeater (Spitler et al)

2018 SEP test with triple system (Archibald et al)

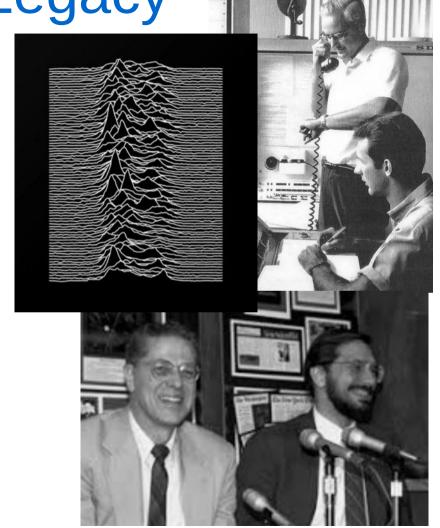
... and much much more (plus non-pulsar science!)





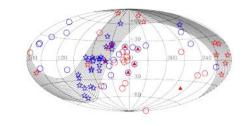








NANOGrav Observing Program





Arecibo Observatory

- Main program: 39 pulsars / dual receiver / every three weeks (began 2004)



Green Bank Telescope

- Main program: 39 pulsars / dual receiver / every month (began 2004)
- High cadence program: 2 pulsars / single receiver /every week (began 2014)



Very Large Array

- Experimental program: 7 pulsars / single or dual receiver
- Sensitive 2-4 GHz system expands frequency coverage
- Can see slightly further south than GBT (1 pulsar)

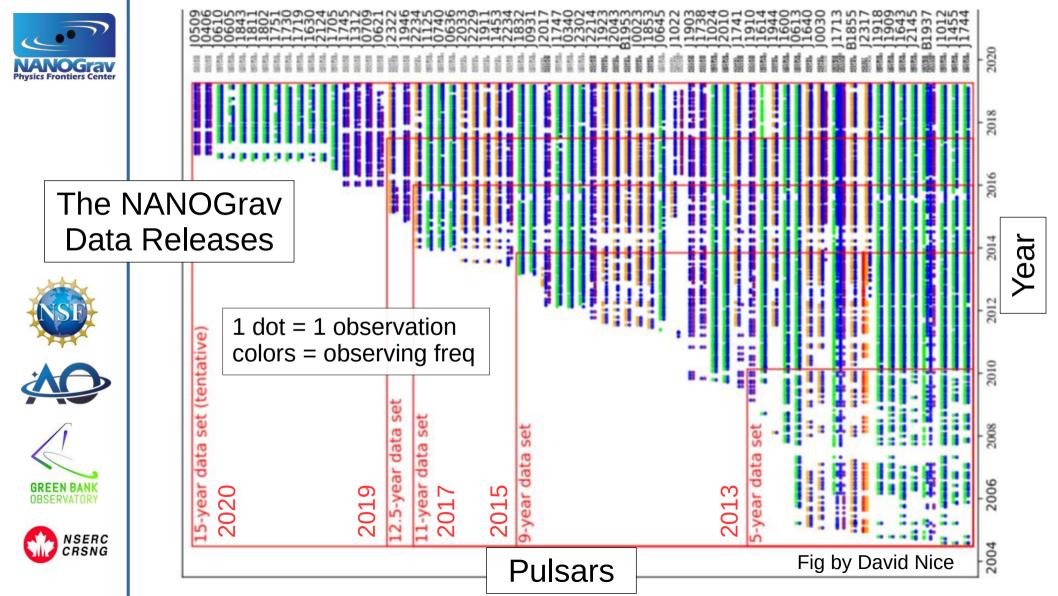


CHIME (with CHIME/Pulsar collaboration)

- 400 to 800 MHz
- All NANOGrav pulsars at $\delta \gtrsim -20^{\circ}$ observed daily (!)
- Dwell time $\sim 5 \min/\cos(\delta)$ (began late 2018)
- Still in engineering phase; challenging to calibrate

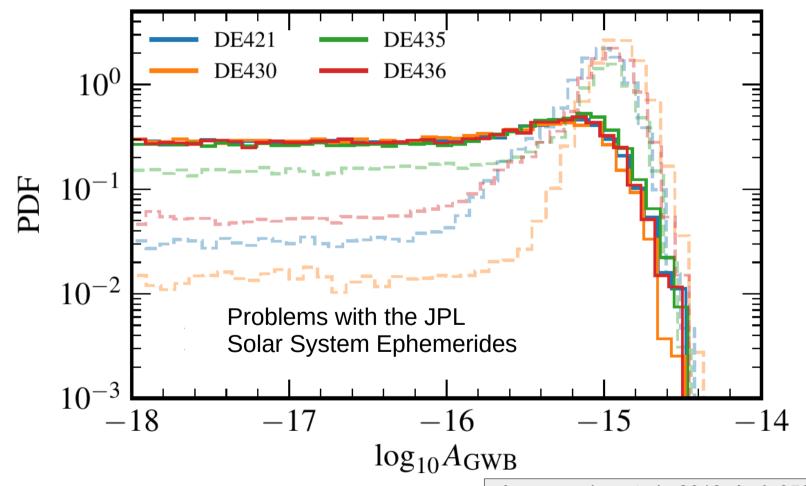


Slide by David Nice





NANOGrav 11 Year GW Bkgd Results











Arzoumanian et al., 2018, ApJ, 859, 47



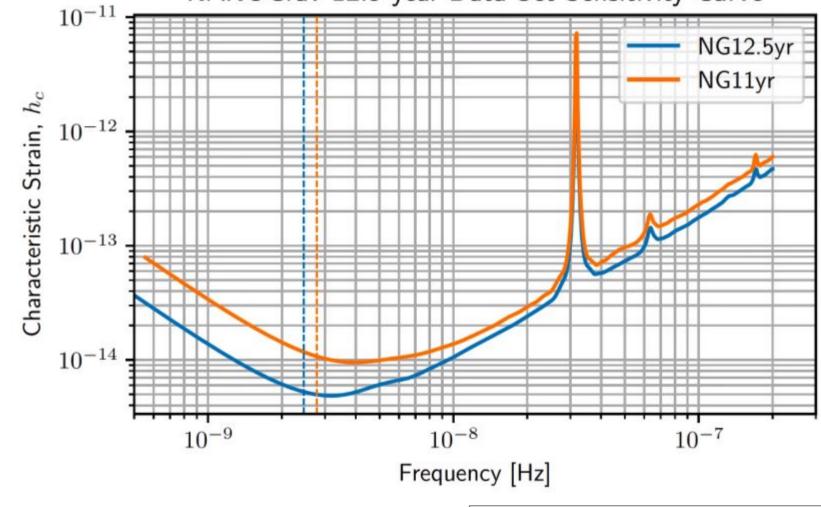










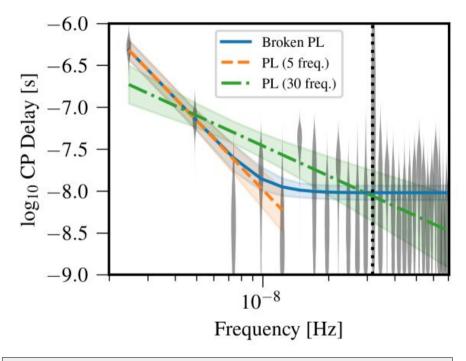


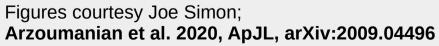
Hazboun, Romano, & Smith, 2019, PhRvD

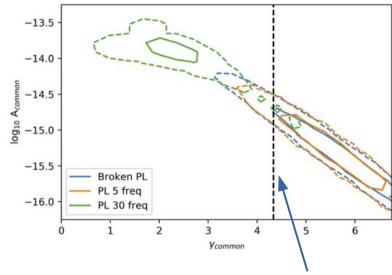


12.5 Year Bkgd Results

Seeing significant low-freq noise in the data....







Expected spectral index for a GWB from inspiraling SMBHBs (γ =13/3)





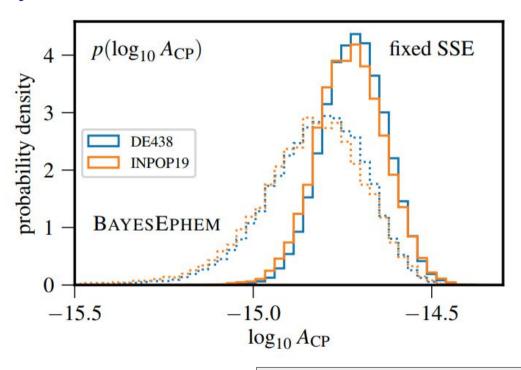


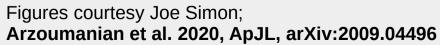




12.5 Year Bkgd Results

Strong evidence for *uncorrelated* common red noise process... (Bayes factors of ~30,000:1 in favor for fixed SS Ephem)











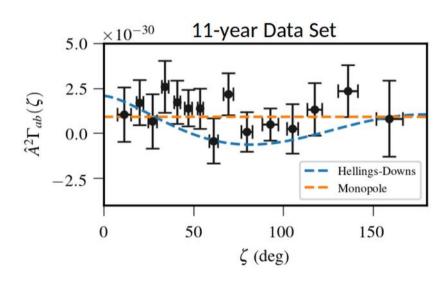


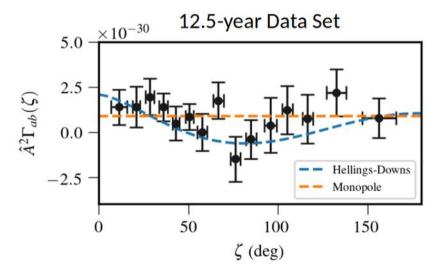


12.5 Year Bkgd Results

No good evidence of the required spatial correlations yet...

But the data are improving rapidly, and there are ~2+ more years of data in the bank!







NSERC CRSNG Figures courtesy Joe Simon;

Arzoumanian et al. 2020, ApJL, arXiv:2009.04496











1. arXiv:2010.04018 [pdf, other] astro-ph.GA

A direct constraint on the Galactic acceleration and the Oort limit from pulsar timing

Authors: Sukanya Chakrabarti, Philip Chang, Michael T, Lam, Sarah J, Vigeland, Alice C, Quillen Submitted 8 October, 2020: originally announced October 2020.

Comments: submitted to Apl Letters

2. arXiv:2010.03976 [pdf, other] asso-ph.co gr-qc hep-th

NANOGrav Hints on Planet-Mass Primordial Black Hol Authors: Guillem Domènech, Shi Pi

Submitted 8 October, 2020; originally announced October 2020.

Comments: 7 pages, 3 figures Report number: IPMU20-0106

3. arXiv:2010.02189 [pdf, other] astro-ph.CO astro-ph.GA hep-ph

Testing Stochastic Gravitational Wave Signals from Pr Authors: Sunao Sugivama, Volodymyr Takhistov, Edoardo Vitagliano Submitted 5 October, 2020: originally announced October 2020.

Comments: 5 pages, 1 figure

Comments: 9 pages, 2 figures

Report number: IPMU20-0105

4. arXiv:2009.14663 [pdf, other] astro-ph.CO gr-qc hep-th

Is the NANOGrav signal a hint of dS decay during infla

Authors: Hao-Hao Li, Gen Ye, Yun-Song Piao Submitted 30 September, 2020; originally announced September 2020.

5. arXiv:2009.14174 [pdf. other] astro-ph.HE astro-ph.HE

NANOGrav signal from MHD turbulence at QCD phase Authors: A. Neronov, A. Roper Pol, C. Caprini, D. Semikoz

Submitted 29 September, 2020; originally announced September 2020. Comments: 5 pages, 2 figures

6. arXiv:2009.13909 [pdf, other] astro-ph.co gr-qc hep-ph Implications of Gravitational-wave Production from D Number of Relativistic Species

Authors: Rvo Namba, Motoo Suzuki

Submitted 5 October, 2020; v1 submitted 29 September, 2020; originally annou Comments: 13 pages, 2 figures, Minor changes, citation additions

7. arXiv:2009.13893 [pdf, other] astro-ph.CO hep-ph

NanoGrav 12.5-yr data and different stochastic Gravit Authors: Ligong Bian, Jing Liu, Ruiyu Zhou

Submitted 29 September, 2020: originally announced September 2020.

Comments: 18 pages, 5 figures, 1 table: comments welcome

8. arXiv:2009.13452 [pdf, other] hep-ph astro-ph.co astro-ph.co astro-ph.co Gravitational wave complementarity and impact of Na Authors: Rome Samanta, Satyabrata Datta

Submitted 30 September, 2020; v1 submitted 28 September, 2020; originally and Comments: 16 pages, 4 figures, typo corrected, refs. updated

9. arXiv:2009.13432 [pdf, other] astro-ph.co gr-qc hep-ph h Implications of the NANOGrav pulsar timing results for

Authors: Sunny Vagnozzi Submitted 28 September, 2020; originally announced September 2020.

Comments: 7 pages, 2 figures, Short spoiler; no. NANOGray is very unlikely to be

10. arXiv:2009.11875 [pdf, other] asso-ph.co hep-ph Whispers from the dark side: Confronting light new pl Authors: Wolfram Ratzinger, Pedro Schwaller

Submitted 24 September, 2020; originally announced September 2020. Comments: 10 pages, 5 figures

Report number: MITP/20-056

11. arXiv:2009.11865 [pdf, other] assro-ph.GA gr-qc

Multimessenger pulsar timing array constraints on supermassive place note pinanes traced by periodic ngine curve

12. arXiv:2009.11853 [pdf, other] astro-ph.co gr-gc hep-ph hep-th

Solar-Mass Primordial Black Holes Explain NANOGrav Hint of Gravitational Waves

Authors: Kazunori Kohri, Takahiro Terada

Next >

GRAVITATIONAL WAVE PULSARS

RANDOM

ASK ME WHAT THE SECRET

WAVES USING PULSARS IS.

TIMING!

TO DETECTING GRAVITATIONAL

WHAT'S THE SECRET

TO DETECTING GRAV-

2020: originally announced September 2020

the title change)

PTC-20-22

hep-ph

table cosmic strings

Kai Schmitz ember 2020

der Phase Transitions

tonino Marciano, Kaigiang Zeng

ber, 2020; originally announced September 2020.

spired GW spectra added, including more recent numerical tools in the subject. New references added, The main conclusions

or-oc hep-ph

ion from Dark Phase Transition: Connecting NANOGrav Pulsar Timing Data and

ou Takahashi, Masaki Yamada

ber, 2020; originally announced September 2020.

the sound-wave period included, figures updated, conclusions unchanged

servations to Monitor Dispersion with the Giant Metrewave Radio Telescope

Jayanta Roy, Michael T. Lam, James M. Cordes, David L. Kaplan, Bhaswati Bhattacharyya, Lina Levin ptember 2020.

hysical Journal

gr-gc hep-ph

Holes as Dark Matter

eptember 2020.

nordial black hole formation?

nber, 2020; originally announced September 2020.

gr-gc hep-ph hep-th for cosmic strings?

iber, 2020; originally announced September 2020.

the higher cosmic-string modes [see Eq. (6)], resulting in a few numerical but no qualitative changes

astro-ph.HE gr-qc hep-ph hep-th

OGrav Pulsar Timing Data

iber, 2020; originally announced September 2020.

royed accuracy of the calculation leading to a minor modification of the results

astro-ph.GA gr-qc

earch For An Isotropic Stochastic Gravitational-Wave Background

rsha Blumer, Bence Becsy, Adam Brazier, Paul R. Brook, Sarah Burke-Spolaor, Shami Chatteriee, Siyuan Chen, vford, H. Thankful Cromartie, Megan E. DeCesar, Paul B. Demorest, Timothy Dolch, Justin A. Ellis, Elizabeth C. an Garver-Daniels, Peter A. Gentile, Deborah C. Good, Jeffrey S. Hazboun, A. Miguel Holgado, et al. (36

stember 2020.

Submitted to The Astrophysical Journal Letters

22. arXiv:2005.13549 [pdf, other] hep-ph astro-ph.co hep-th

Gravitational waves and proton decay: complementary windows into GUTs Authors: Stephen F. King, Silvia Pascoli, lessica Turner, Ye-Ling Zhou

Authors: Chengcheng Xin, Chiara M. F. Mingarelli, Jeffrey S. Hazboun Submitted 24 September, 2020; originally announced September 2020. Comments: 11 pages, 4 figures, submitted to ApJ



Predictions for Future

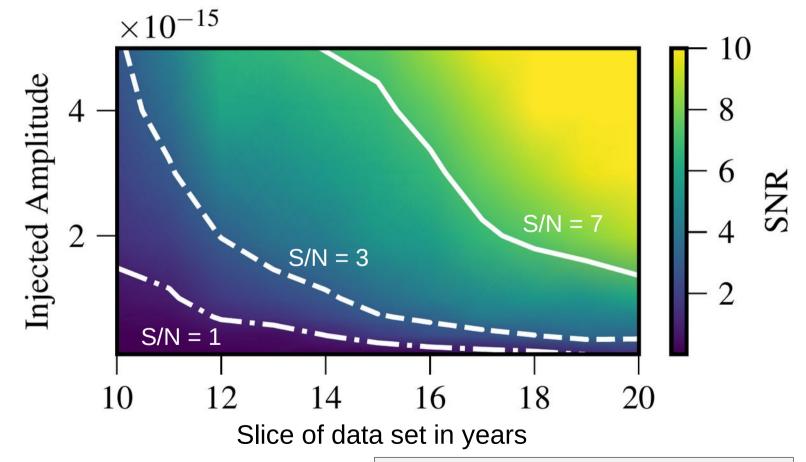


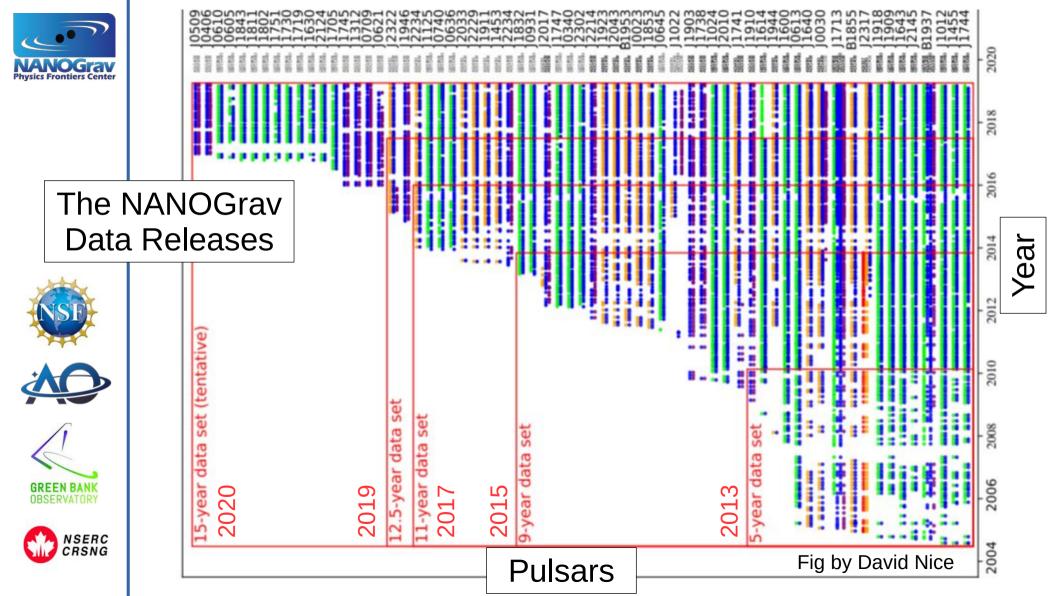


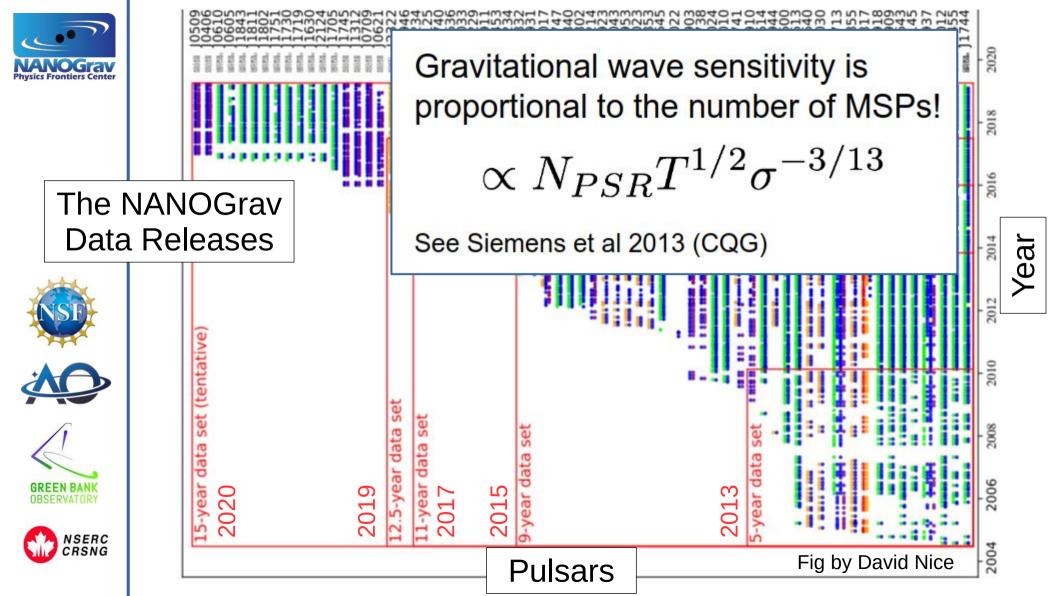






Fig courtesy of Nihan Pol (arXiv:2010.11950)







Predictions for Future

Predictions for future MSPs, using AO, GBT, and DSA-2000 telescopes

~200 MSPs timeable at ~1µs

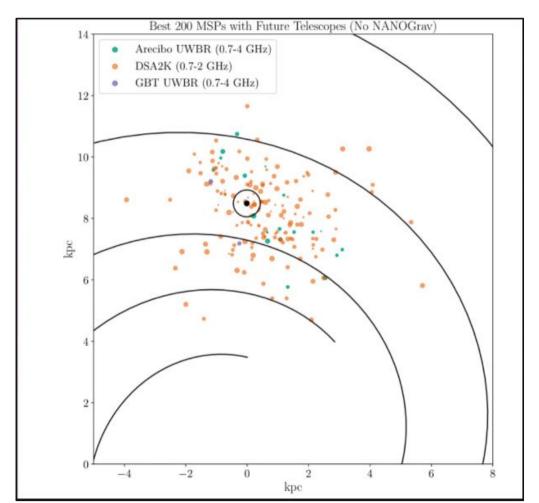
Fig courtesy of Tyler Coher and Paul Demorest











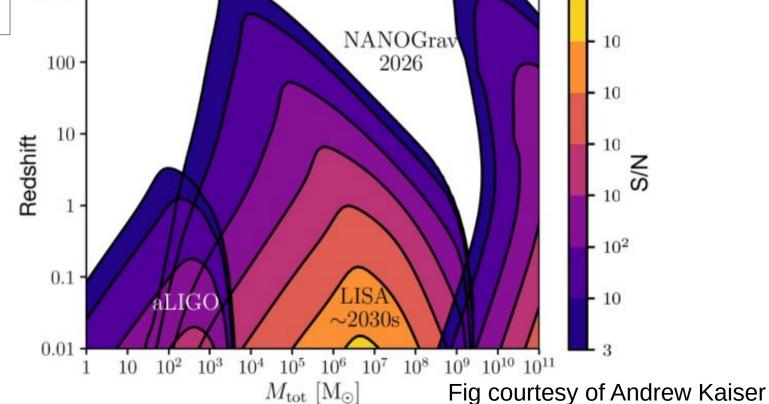


Predictions for Future

Should detect one or more individual SMBHBs by the end of the decade

Individual SMBHB Sources

1000 - 1000













Wideband Receiver for GBT

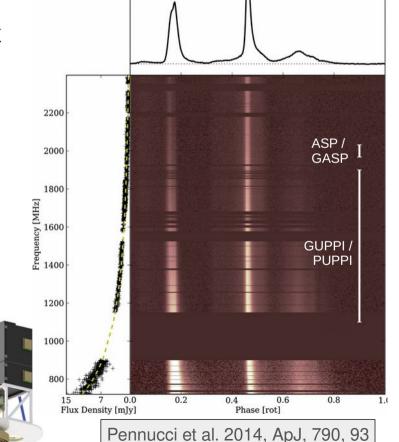
- Need multi-frequencies for ISM removal
 - Current systems have <800 MHz BW
 - Ultra-wideband system would give 2x better timing, fewer systematics, and more protection from scintillation
- Building a 0.7-4 GHz receiver for GBT, funded by Moore Foundation





Hobbs et al. 2019, PASA, in press















Summary

- NANOGrav is doing great and we are optimistic
- Arecibo loss is terrible. Will slow, not stop progress.
 - In long term, we need a replacement (e.g. DSA-2000)
- We have 15+ years of data in hand
- Our work with IPTA will make things even better
- Data are intriguing and we expect a detection within the next couple years
- In the meantime, tons of other science (e.g. Cromartie et al. 2019 massive NS)







